# **Graphical Models**

Data 102 Fall 2022 Lecture 7

## Weekly Overview

• So far:

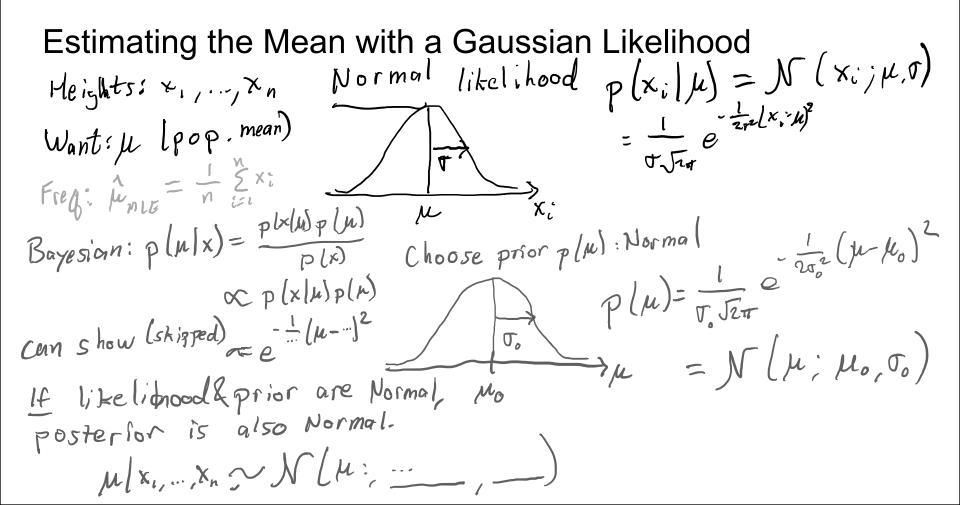
### • Today:

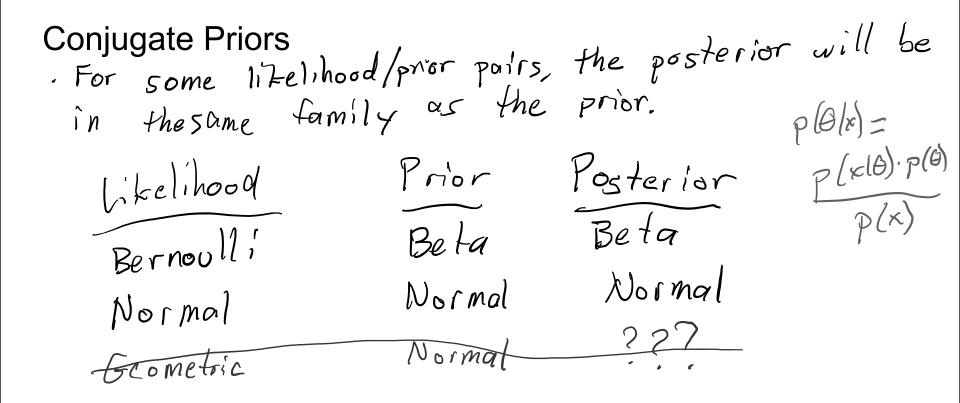
- Making decisions with feedback (online decision-making)
- Hypothesis testing with a known alternative (Neyman-Pearson)
- Connection to binary classification

• Next time: connecting decision-making and frequentist/Bayesian views

## Recap: Statistical modeling

- Goal: find unknown parameter  $\theta$  using observed data x
- How:
  - Define a probability model for the data/parameters, then use it to estimate  $\theta$  from x
- Likelihood function p(x|θ): captures how likely our data are for each parameter
  - Used in both frequentist and Bayesian models
- Frequentist modeling
  - MLE (Maximum Likelihood Estimate): value of  $\theta$  that makes  $p(x|\theta)$  as large as possible
- Bayesian modeling
  - Define a prior  $p(\theta)$ : what we believe about the parameter before we see any data
  - $\circ$  Compute posterior p( $\theta|x)$ : what we believe about the parameter after observing data
  - To get a single estimate for  $\theta$  from the posterior, we can use the MAP or LMSE estimates
  - MAP: value of  $\theta$  that makes  $p(\theta|x)$  as large as possible
  - *MMSE*:  $E_{\theta|x}[\theta]$  (expectation of  $\theta$  according to posterior  $p(\theta|x)$ )





Assumptions · Planets are cid (given grp mens) Exoplanet Model · Jo, J, fixed · Each group's radii nor mally det. Mave: X.,..., Xn (radii of exoplanets) Prior mouns come from same dist. Lychoose Op large  $z_i$  is 20 o.w. Warts · Mean raulius for each "group" · Which planets are in each grp? · fixed Zi : is planet i habitable?  $Z_{i} \stackrel{\text{id}}{\xrightarrow{}} Bernoulli(\pi)$   $\mu_{0} \sim \mathcal{N}(\mu_{p}, \tau_{p})$   $\mu_{1} \sim \mathcal{N}(\mu_{p}, \tau_{p})$   $\mu_{1} \sim \mathcal{N}(\mu_{p}, \tau_{p})$ Mo , To : mean /SD for "habitable" Mo , To : " for "gas giant" Posterior\_  $\chi_{i}|_{z_{i},\mu_{o},\mu_{i}} \sim \mathcal{N}(\mu_{z_{i}}, \tau_{z_{i}})$ P(Z.,...,Zn,Mo, M, (X.,...,Xn) Glikelihood  $\Theta = [z_1, \dots, z_n, \mu_0, \mu)$ 

#### **Exoplanet Model: A Visual Representation** Π $z_i \sim \text{Bernoulli}(\pi)$ $\mu_k \sim \mathcal{N}(\mu_p, \sigma_p) \; k = 0, k = 1$ 2n $z_1^{'}$ Zi $x_i | z_i, \mu_0, \mu_1 \sim \mathcal{N}(\mu_{z_i}, \sigma)$ Graphical Models x<sub>n</sub> . Circles for RUS . Dots for params $\mu_l$ μo J.J. . Arrows indicate dependence

MP, Jp